

The Regularity Transformation Equations:

An Elliptic Mechanism for Smoothing Gravitational Metrics in General Relativity

Blake Temple
Applied Math Seminar, UC-Davis
October 25, 2018

*All Joint Work With: **Moritz Reintjes***

- I discuss recent work with Moritz Reintjes in which we derive the Regularity Transformation equations, a system of nonlinear elliptic equations with matrix valued differential forms as unknowns.

The RT-equations:

$$\Delta \tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1}dJ) + d(J^{-1}A), \quad (1)$$

$$\Delta J = \delta(J\Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A, \quad (2)$$

$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle}), \quad (3)$$

$$\delta \vec{A} = v, \quad (4)$$

$$\text{Curl}(J) \equiv \partial_j J_i^\mu - \partial_i J_j^\mu = 0 \quad \text{on } \partial\Omega, \quad (5)$$

- We prove that existence of solutions to the RT-equations is equivalent to the existence of a coordinate transformation sufficient to smooth a crinkled map of spacetime to optimal metric regularity in General Relativity.

- We then give an **existence theorem** for the RT-equations **based on elliptic regularity** in L^p –spaces...the result:

If a connection and its curvature tensor are both in $W^{m,p}$, $m \geq 1, p > n$, then there always exists a coordinate transformation with Jacobian in $W^{m+1,p}$, such that in the new coordinate system, the connection is in $W^{m+1,p}$

- This tells us that we can solve the Einstein equations in coordinate systems in which the equations are simpler and solution metrics are one order below optimal, and still be guaranteed the existence of other coordinate systems in which the metric exhibits optimal regularity, i.e., two derivatives above its curvature tensor.

- When connection and curvature are in L^∞ , the RT-equations reduce the open problem of regularity singularities at GR shock waves to an approachable problem in elliptic regularity theory, a topic of authors' current research.

- The starting point for the derivation of the RT-equations is the Riemann-flat condition, a geometric condition for metric smoothing introduced previously by the authors.

Introduction

Thesis Statement:

The Einstein equations of General Relativity are tensorial—Given independent of coordinates...

$$G = \kappa T$$

To solve them, one has to specify a coordinate system in which the equations take a form simple enough to solve by known methods of **PDE's**...

The spacetime metric may not exhibit its optimal level of regularity in the coordinates in which the equations can be solved.

In Einstein's theory of General Relativity:

The Einstein equations $G = \kappa T$

are equns for the gravitational metric $g = g_{ij}$

coupled to the fluid sources ρ, p, u

$$G_{ij}[g_{ij}] = \kappa T_{ij}(\rho, p, u)$$

$$Div T = T_{j;\sigma}^{\sigma} = 0$$

In Einstein's theory of General Relativity:

The gravitational metric tensor g determines the properties of spacetime...

geodesics, parallel translation, time dilation, arc length...
...as well as the connection Γ and curvature R

Connection:
$$\Gamma_{jk}^i = \frac{1}{2} g^{i\sigma} \{-g_{jk,i} + g_{i,jk} + g_{ki,j}\}$$

Riemann
Curvature:
$$R_{ijk}^l = \Gamma_{ik,j}^l - \Gamma_{ij,k}^l + \Gamma_{j\sigma}^l \Gamma_{ik}^\sigma - \Gamma_{k\sigma}^l \Gamma_{ij}^\sigma$$

Our Question: Given a non-optimal solution of the Einstein equations, does there always exist a coordinate transformation which smooths the spacetime to its Optimal Regularity??

Our Question: Given a non-optimal solution of the Einstein equations, does there always exist a coordinate transformation which smooths the spacetime to its Optimal Regularity??

By Optimal Regularity, we mean that the gravitational metric tensor is two full derivatives smoother than its curvature tensor.

Our Question: Given a non-optimal solution of the Einstein equations, does there always exist a coordinate transformation which smooths the spacetime to its Optimal Regularity??

By Optimal Regularity, we mean that the gravitational metric tensor is two full derivatives smoother than its curvature tensor.

The equations may be too complicated to solve in coordinates in which the metric exhibits its optimal regularity...

Motivation: Non-optimal solutions of the Einstein equations first came up when we constructed shock wave solutions of the Einstein equations by the Glimm Scheme in Standard Schwarzschild Coordinates (SSC)

Groah and I observed: ``The gravitational metric appears to be singular at the shocks in coordinates where the analysis is feasible (SSC)’’.

The Einstein Equations In SSC

Assume a gravitational metric ansatz of the SSC form:

$$ds^2 = -B(t, r)dt^2 + \frac{dr^2}{A(t, r)} + r^2 d\Omega^2$$

Plug into the Einstein equations :

$$G = \kappa T$$

$$T_{ij} = (\rho + p)u_i u_j + p g_{ij}$$

Standard Schwarzschild Coordinates

Four
PDE's

$$\left\{ -r \frac{A_r}{A} + \frac{1-A}{A} \right\} = \frac{\kappa B}{A} r^2 T^{00} \quad (1)$$

$$\frac{A_t}{A} = \frac{\kappa B}{A} r T^{01} \quad (2)$$

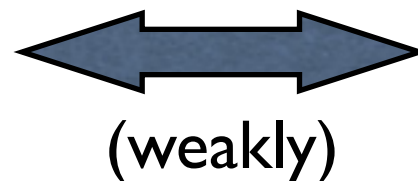
$$\left\{ r \frac{B_r}{B} - \frac{1-A}{A} \right\} = \frac{\kappa}{A^2} r^2 T^{11} \quad (3)$$

$$- \left\{ \left(\frac{1}{A} \right)_{tt} - B_{rr} + \Phi \right\} = 2 \frac{\kappa B}{A} r^2 T^{22}, \quad (4)$$

where

$$\begin{aligned} \Phi = & \frac{B_t A_t}{2A^2 B} - \frac{1}{2A} \left(\frac{A_t}{A} \right)^2 - \frac{B_r}{r} - \frac{B A_r}{r A} \\ & + \frac{B}{2} \left(\frac{B_r}{B} \right)^2 - \frac{B}{2} \frac{B_r}{B} \frac{A_r}{A}. \end{aligned}$$

(1)+(2)+(3)+(4)



(1)+(3)+div T=0

Theorem: (Te-Gr) The equations close in a “locally inertial” formulation of (1), (2) & Div T=0:

$$\{T_M^{00}\}_{,0} + \left\{ \sqrt{AB} T_M^{01} \right\}_{,1} = -\frac{2}{r} \sqrt{AB} T_M^{01}, \quad (1)$$

$$\begin{aligned} \{T_M^{01}\}_{,0} + \left\{ \sqrt{AB} T_M^{11} \right\}_{,1} = & -\frac{1}{2} \sqrt{AB} \left\{ \frac{4}{r} T_M^{11} + \frac{(1-A)}{Ar} (T_M^{00} - T_M^{11}) \right. \\ & \left. + \frac{2\kappa r}{A} (T_M^{00} T_M^{11} - (T_M^{01})^2) - 4r T^{22} \right\}, \end{aligned} \quad (2)$$

$$r A_r = (1-A) - \kappa r^2 T_M^{00}, \quad (3)$$

$$r B_r = \frac{B(1-A)}{A} + \frac{B}{A} \kappa r^2 T_M^{11}. \quad (4)$$

$$T_M^{00} = \frac{\rho c^2 + p}{1 - \left(\frac{v}{c}\right)^2} \quad T_M^{01} = \frac{\rho c^2 + p}{1 - \left(\frac{v}{c}\right)^2} \frac{v}{c}$$

$$T_M^{11} = \frac{p + \left(\frac{v}{c}\right)^2}{1 - \left(\frac{v}{c}\right)^2} \rho c^2 \quad T^{22} = \frac{p}{r^2} \quad v = \frac{1}{\sqrt{AB}} \frac{u^1}{u^0}$$

$$\{T_M^{00}\}_{,0} + \left\{ \sqrt{AB} T_M^{01} \right\}_{,1} = -\frac{2}{r} \sqrt{AB} T_M^{01}, \quad (1)$$

$$\begin{aligned} \{T_M^{01}\}_{,0} + \left\{ \sqrt{AB} T_M^{11} \right\}_{,1} = & -\frac{1}{2} \sqrt{AB} \left\{ \frac{4}{r} T_M^{11} + \frac{(1-A)}{Ar} (T_M^{00} - T_M^{11}) \right. \\ & \left. + \frac{2\kappa r}{A} (T_M^{00} T_M^{11} - (T_M^{01})^2) - 4r T^{22} \right\}, \end{aligned} \quad (2)$$

$$r A_r = (1-A) - \kappa r^2 T_M^{00}, \quad (3)$$

$$r B_r = \frac{B(1-A)}{A} + \frac{B}{A} \kappa r^2 T_M^{11}. \quad (4)$$

$$\{T_M^{00}\}_{,0} + \left\{ \sqrt{AB} T_M^{01} \right\}_{,1} = -\frac{2}{r} \sqrt{AB} T_M^{01}, \quad (1)$$

$$\begin{aligned} \{T_M^{01}\}_{,0} + \left\{ \sqrt{AB} T_M^{11} \right\}_{,1} = & -\frac{1}{2} \sqrt{AB} \left\{ \frac{4}{r} T_M^{11} + \frac{(1-A)}{Ar} (T_M^{00} - T_M^{11}) \right. \\ & \left. + \frac{2\kappa r}{A} (T_M^{00} T_M^{11} - (T_M^{01})^2) - 4r T^{22} \right\}, \end{aligned} \quad (2)$$

$$r A_r = (1-A) - \kappa r^2 T_M^{00}, \quad (3)$$

$$r B_r = \frac{B(1-A)}{A} + \frac{B}{A} \kappa r^2 T_M^{11}. \quad (4)$$



The metric components A,B...

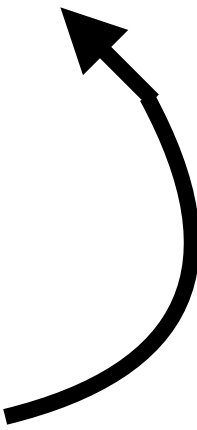
$$\{T_M^{00}\}_{,0} + \left\{ \sqrt{AB} T_M^{01} \right\}_{,1} = -\frac{2}{r} \sqrt{AB} T_M^{01}, \quad (1)$$

$$\begin{aligned} \{T_M^{01}\}_{,0} + \left\{ \sqrt{AB} T_M^{11} \right\}_{,1} = & -\frac{1}{2} \sqrt{AB} \left\{ \frac{4}{r} T_M^{11} + \frac{(1-A)}{Ar} (T_M^{00} - T_M^{11}) \right. \\ & \left. + \frac{2\kappa r}{A} (T_M^{00} T_M^{11} - (T_M^{01})^2) - 4r T^{22} \right\}, \end{aligned} \quad (2)$$

$$r A_r = (1-A) - \kappa r^2 T_M^{00}, \quad (3)$$

$$r B_r = \frac{B(1-A)}{A} + \frac{B}{A} \kappa r^2 T_M^{11}. \quad (4)$$

The metric components A,B...
 ...are only one derivative
 smoother than the sources T



$$\{T_M^{00}\}_{,0} + \left\{ \sqrt{AB} T_M^{01} \right\}_{,1} = -\frac{2}{r} \sqrt{AB} T_M^{01}, \quad (1)$$

$$\begin{aligned} \{T_M^{01}\}_{,0} + \left\{ \sqrt{AB} T_M^{11} \right\}_{,1} = & -\frac{1}{2} \sqrt{AB} \left\{ \frac{4}{r} T_M^{11} + \frac{(1-A)}{Ar} (T_M^{00} - T_M^{11}) \right. \\ & \left. + \frac{2\kappa r}{A} (T_M^{00} T_M^{11} - (T_M^{01})^2) - 4r T^{22} \right\}, \end{aligned} \quad (2)$$

$$r A_r = (1-A) - \kappa r^2 T_M^{00}, \quad (3)$$

$$r B_r = \frac{B(1-A)}{A} + \frac{B}{A} \kappa r^2 T_M^{11}. \quad (4)$$

The metric components A,B...
 ...are only one derivative
 smoother than the sources T

Since $G = \kappa T \rightarrow$

$$\{T_M^{00}\}_{,0} + \left\{ \sqrt{AB} T_M^{01} \right\}_{,1} = -\frac{2}{r} \sqrt{AB} T_M^{01}, \quad (1)$$

$$\begin{aligned} \{T_M^{01}\}_{,0} + \left\{ \sqrt{AB} T_M^{11} \right\}_{,1} = & -\frac{1}{2} \sqrt{AB} \left\{ \frac{4}{r} T_M^{11} + \frac{(1-A)}{Ar} (T_M^{00} - T_M^{11}) \right. \\ & \left. + \frac{2\kappa r}{A} (T_M^{00} T_M^{11} - (T_M^{01})^2) - 4r T^{22} \right\}, \end{aligned} \quad (2)$$

$$r A_r = (1-A) - \kappa r^2 T_M^{00}, \quad (3)$$

$$r B_r = \frac{B(1-A)}{A} + \frac{B}{A} \kappa r^2 T_M^{11}. \quad (4)$$

The metric components A,B...
...are only one derivative
smoother than the sources T

Since $G = \kappa T \rightarrow$

The metric is only one order
derivative smoother than
the curvature tensor...

Conclude: For shock wave solutions of the Einstein equations $G = \kappa T$ generated by the Glimm Scheme:

Conclude: For shock wave solutions of the Einstein equations $G = \kappa T$ generated by the Glimm Scheme:

● Solution:
$$ds^2 = -B(t, r)dt^2 + \frac{dr^2}{A(t, r)} + r^2 d\Omega^2$$

A, B are Lipschitz Continuous

Conclude: For shock wave solutions of the Einstein equations $G = \kappa T$ generated by the Glimm Scheme:

● Solution:
$$ds^2 = -B(t, r)dt^2 + \frac{dr^2}{A(t, r)} + r^2 d\Omega^2$$

A, B are Lipschitz Continuous

● Riemann Curvature Tensor: $R^i_{l k j}(t, r) \in L^\infty$

Conclude: For shock wave solutions of the Einstein equations $G = \kappa T$ generated by the Glimm Scheme:

● Solution:
$$ds^2 = -B(t, r)dt^2 + \frac{dr^2}{A(t, r)} + r^2 d\Omega^2$$

A, B are Lipschitz Continuous

● Riemann Curvature Tensor: $R^i_{l k j}(t, r) \in L^\infty$

● Fluid variables: $\rho(t, r), p(r, t), v(r, t) \in L^\infty$

Conclude: For shock wave solutions of the Einstein equations $G = \kappa T$ generated by the Glimm Scheme:

● Solution:
$$ds^2 = -B(t, r)dt^2 + \frac{dr^2}{A(t, r)} + r^2 d\Omega^2$$

A, B are Lipschitz Continuous

● Riemann Curvature Tensor: $R^i_{l k j}(t, r) \in L^\infty$

● Fluid variables: $\rho(t, r), p(r, t), v(r, t) \in L^\infty$

Conclude: Second derivatives of the metric contain delta function sources, but these cancel out in the curvature tensor...

This is a most natural setting for shock waves in GR because the Einstein equations $G = \kappa T$ place $G \in L^\infty$ when the bounded discontinuous fluid sources are $T \in L^\infty$.

In General: Given a Metric g_{ij} ...

In General: Given a Metric $g_{ij} \dots$

Connection:
$$\Gamma_{jk}^i = \frac{1}{2} g^{i\sigma} \{ -g_{jk,i} + g_{i,jk} + g_{ki,j} \}$$

In General: **Given a Metric** $g_{ij} \dots$

Connection:
$$\Gamma_{jk}^i = \frac{1}{2} g^{i\sigma} \{-g_{jk,i} + g_{i,jk} + g_{ki,j}\}$$

**Riemann
Curvature:**
$$R_{ijk}^l = \Gamma_{ik,j}^l - \Gamma_{ij,k}^l + \Gamma_{j\sigma}^l \Gamma_{ik}^\sigma - \Gamma_{k\sigma}^l \Gamma_{ij}^\sigma$$

In General: Given a Metric $g_{ij} \dots$

Connection:
$$\Gamma_{jk}^i = \frac{1}{2} g^{i\sigma} \{ -g_{jk,i} + g_{i,jk} + g_{ki,j} \}$$

Riemann Curvature:
$$R_{ijk}^l = \Gamma_{ik,j}^l - \Gamma_{ij,k}^l + \Gamma_{j\sigma}^l \Gamma_{ik}^\sigma - \Gamma_{k\sigma}^l \Gamma_{ij}^\sigma$$

“R is a curl plus a commutator”

$$R \approx d\Gamma + [\Gamma, \Gamma]$$

In General: **Given a Metric** $g_{ij} \dots$

Connection: $\Gamma_{jk}^i = \frac{1}{2} g^{i\sigma} \{-g_{jk,i} + g_{i,jk} + g_{ki,j}\}$

Riemann Curvature: $R_{ijk}^l = \Gamma_{ik,j}^l - \Gamma_{ij,k}^l + \Gamma_{j\sigma}^l \Gamma_{ik}^\sigma - \Gamma_{k\sigma}^l \Gamma_{ij}^\sigma$

“R is a curl **plus** a commutator” $R \approx d\Gamma + [\Gamma, \Gamma]$

Conclude:

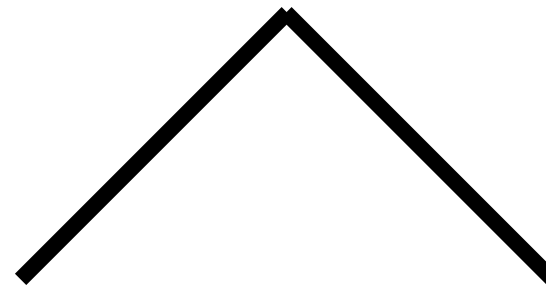
$$\Gamma \approx \partial g$$

$$R \approx \partial^2 g$$

Thus for Shock-Waves:

Thus for Shock-Waves:

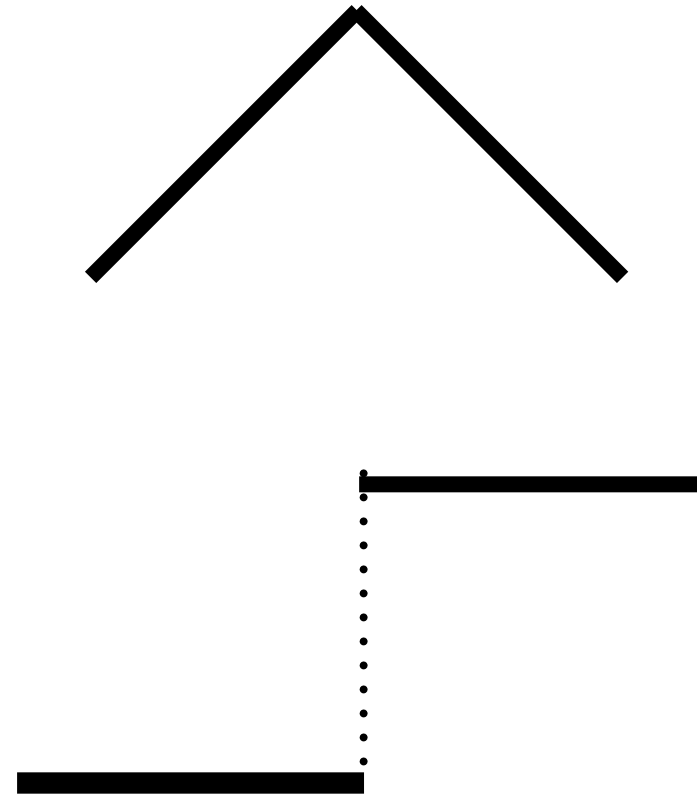
g is Lipschitz



Thus for Shock-Waves:

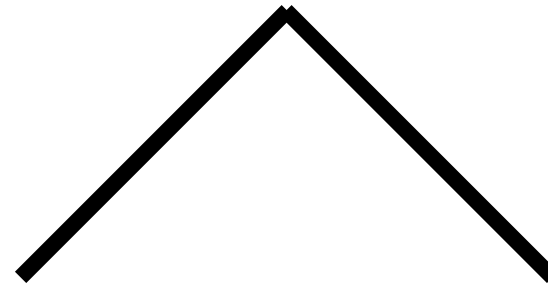
g is Lipschitz

$\Gamma \approx \partial g$ is L^∞

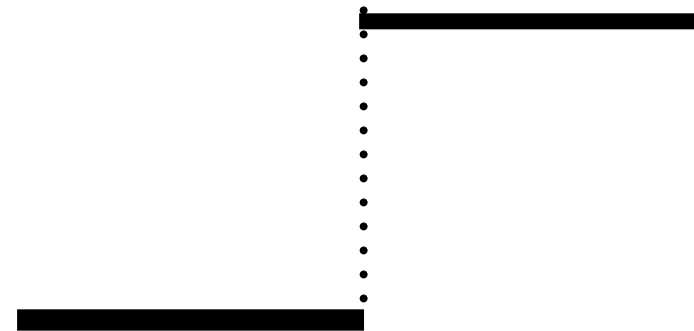


Thus for Shock-Waves something is special:

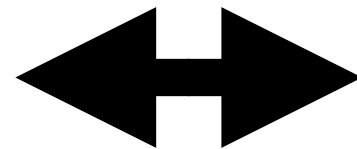
g is Lipschitz



$\Gamma \approx \partial g$ is L^∞



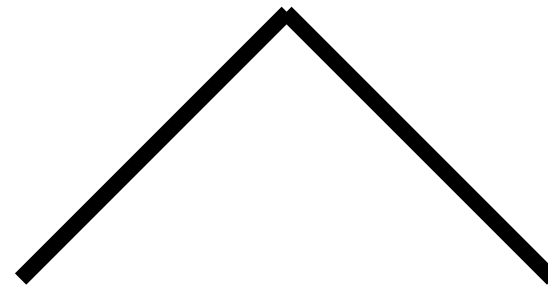
$R \approx \partial^2 g$ is L^∞



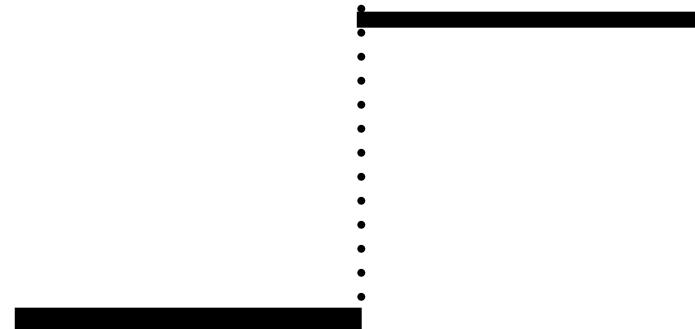
$d\Gamma \in L^\infty$

Thus for Shock-Waves something is special:

g is Lipschitz



$\Gamma \approx \partial g$ is L^∞

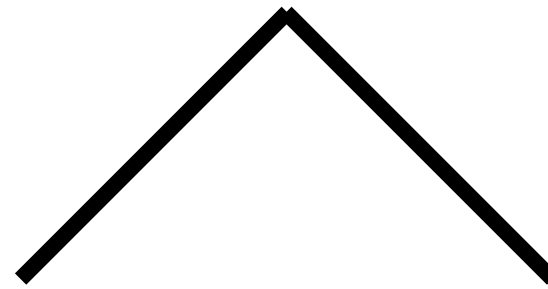


$$R \approx \partial^2 g \text{ is } L^\infty \quad \longleftrightarrow \quad d\Gamma \in L^\infty$$

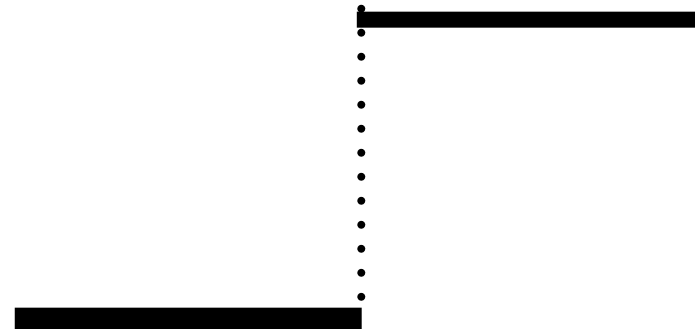
In SSC... The curvature is only one derivative less smooth than g ...

Thus for Shock-Waves something is special:

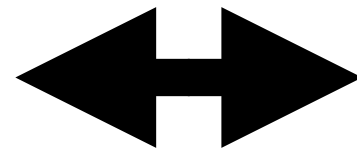
g is Lipschitz



$\Gamma \approx \partial g$ is L^∞



$R \approx \partial^2 g$ is L^∞



$d\Gamma \in L^\infty$

So all delta functions cancel out in $d\Gamma$

Conclude: For shock waves we have....

$$\Gamma, d\Gamma \in L^\infty \quad \text{so} \quad g \in C^{0,1}$$

...in x -coordinates

We ask: “Does there exist a coordinate transformation $x \rightarrow y$ such that in y -coordinates, we have...”

$$\Gamma \in C^{0,1} \quad \text{so} \quad g \in C^{1,1} \quad ?$$

$$(d\Gamma, R \in L^\infty)$$

More Generally: **Given**....

$$\Gamma, d\Gamma \in W^{m,p} \quad \text{so} \quad g \in W^{m+1,p}$$

“Does there exist a coordinate transformation $x \rightarrow y$ such that in y -coordinates...”

$$\Gamma \in W^{m+1,p} \quad \text{so} \quad g \in W^{m+2,p} \quad ?$$

$$(d\Gamma, R \in W^{m,p})$$

In words: If the connection has the same regularity as the curvature in some coordinate system, does there always exist a coordinate transformation which smooths the connection, (and hence the metric), by one order ??

Defn: We say that a connection has “optimal regularity” if it is two orders more regular than its curvature tensor.

Defn: We say that a connection has “optimal regularity” if it is two orders more regular than its curvature tensor.

Our Question: Can “non-optimal” connections always be lifted to optimal regularity by coordinate transformation?

Defn: If Γ non-optimal and no such coordinate transformation exists at P , we call P a “Regularity Singularity”

Defn: If Γ non-optimal and no such coordinate transformation exists at P , we call P a “Regularity Singularity”

For example: If P is a regularity singularity on a shock wave, then spacetime is not regular enough to admit locally inertial coordinate frames:

$$g \in C^{0,1} \text{ (Lipschitz continuous)}$$

$$g_{ij}(P) = \eta \equiv \text{diag}(-1, 1, 1, 1), \quad g_{ij,k}(P) = 0 \quad ???$$

Conclude: If shock-wave interaction can create a regularity singularity, then spacetime is not locally Minkowski...

I.e., the physics of GR does not reduce to Special Relativity...

New GR scattering effects would occur in a neighborhood of a shock-wave Regularity Singularity

.....

Regularity Singularities and the scattering of gravity waves in approximate locally inertial frames,

M. Reintjes and B. Temple, Meth Appl Anal, Vol. 23, No. 2, pp. 233-258, September 2016.

Our Question: Given

$$\Gamma, d\Gamma \in W^{m,p} \quad \text{so} \quad g \in W^{m+1,p}$$

“Does there exist a coordinate transformation $x \rightarrow y$ such that in y -coordinates...”

$$\Gamma \in W^{m+1,p} \quad \text{so} \quad g \in W^{m+2,p} \quad ?$$

$$(d\Gamma, R \in W^{m,p})$$

Our Question: Given

$$\Gamma, d\Gamma \in W^{m,p} \quad \text{so} \quad g \in W^{m+1,p}$$

“Does there exist a coordinate transformation $x \rightarrow y$ such that in y -coordinates...”

$$\Gamma \in W^{m+1,p} \quad \text{so} \quad g \in W^{m+2,p} \quad ?$$

$$(d\Gamma, R \in W^{m,p})$$

Question: How smooth should the Jacobian be?

Answer: The Jacobian should have the same regularity as the metric...

$$J \equiv \frac{\partial x^i}{\partial y^\mu} \in W^{m+1,p}$$

That is: ...

$$\bar{g}_{\mu\nu} = \frac{\partial x^i}{\partial y^\mu} g_{ij} \frac{\partial x^j}{\partial y^\nu}$$

The diagram illustrates the regularity of the components in the equation $\bar{g}_{\mu\nu} = \frac{\partial x^i}{\partial y^\mu} g_{ij} \frac{\partial x^j}{\partial y^\nu}$. Arrows point from the regularity spaces to the corresponding terms:

- $W^{m+2,p}$ points to $\bar{g}_{\mu\nu}$
- $W^{m+1,p}$ points to $\frac{\partial x^i}{\partial y^\mu}$
- $W^{m+1,p}$ points to g_{ij}
- $W^{m+1,p}$ points to $\frac{\partial x^j}{\partial y^\nu}$

For shock-waves: $g \in C^{0,1}, \Gamma \in L^\infty, J \in C^{0,1}$

That is:

$$\bar{g}_{\mu\nu} = \frac{\partial x^i}{\partial y^\mu} g_{ij} \frac{\partial x^j}{\partial y^\nu}$$

$C^{1,1}$ $C^{0,1}$ $C^{0,1}$ $C^{0,1}$

“We need discontinuities in derivatives to cancel out in the Leibniz products...”

For shock-waves: $g \in C^{0,1}$, $\Gamma \in L^\infty$, $J \in C^{0,1}$

That is:

$$\bar{g}_{\mu\nu} = \frac{\partial x^i}{\partial y^\mu} g_{ij} \frac{\partial x^j}{\partial y^\nu}$$

The diagram illustrates the regularity requirements for the components of the metric transformation formula. Arrows point from the regularity classes to the corresponding terms in the equation:

- An arrow from $C^{1,1}$ points to $\bar{g}_{\mu\nu}$.
- An arrow from $C^{0,1}$ points to $\frac{\partial x^i}{\partial y^\mu}$.
- An arrow from $C^{0,1}$ points to g_{ij} .
- An arrow from $C^{0,1}$ points to $\frac{\partial x^j}{\partial y^\nu}$.

“To smooth out a metric singularity requires a singular transformation...”

The connection
at shock-waves:

$$g \in C^{0,1}, \quad \Gamma \in L^\infty, \quad J \in C^{0,1}$$

That is:

$$\tilde{\Gamma}_{\beta\gamma}^\alpha = \Gamma_{jk}^i \frac{\partial y^\alpha}{\partial x^i} \frac{\partial x^j}{\partial y^\beta} \frac{\partial x^k}{\partial y^\gamma} + \frac{\partial^2 y^\alpha}{\partial x^\sigma \partial x^\tau} \frac{\partial x^\sigma}{\partial y^\beta} \frac{\partial x^\tau}{\partial y^\gamma}$$

$C^{0,1}$ L^∞ $C^{0,1}$ L^∞ $C^{0,1}$

“Discontinuities have to cancel out on the RHS
to smooth the connection...”

The curvature
at shock-waves:

$$g \in C^{0,1}, \Gamma \in L^\infty, J \in C^{0,1}$$

That is:

$$\tilde{R}^\alpha_{\beta\gamma\delta} = R^i_{jkl} \frac{\partial y^\alpha}{\partial x^i} \frac{\partial x^j}{\partial y^\beta} \frac{\partial x^k}{\partial y^\gamma} \frac{\partial x^l}{\partial y^\delta}$$

The diagram illustrates the regularity requirements for the curvature tensor. Two arrows labeled L^∞ point to the tensors $\tilde{R}^\alpha_{\beta\gamma\delta}$ and R^i_{jkl} in the equation. Four arrows labeled $C^{0,1}$ point to the partial derivative terms $\frac{\partial y^\alpha}{\partial x^i}$, $\frac{\partial x^j}{\partial y^\beta}$, $\frac{\partial x^k}{\partial y^\gamma}$, and $\frac{\partial x^l}{\partial y^\delta}$.

“The curvature involves 2nd derivatives, but it is a tensor, so J maintains the regularity of the curvature R in $L^\infty \dots$ ”

Theorem (SmTe): $C^{1,1}$ coordinate transformations with $C^{0,1}$ Jacobians preserve the weak formulation of the Einstein equations $G = \kappa T$ at shocks

Our Question:

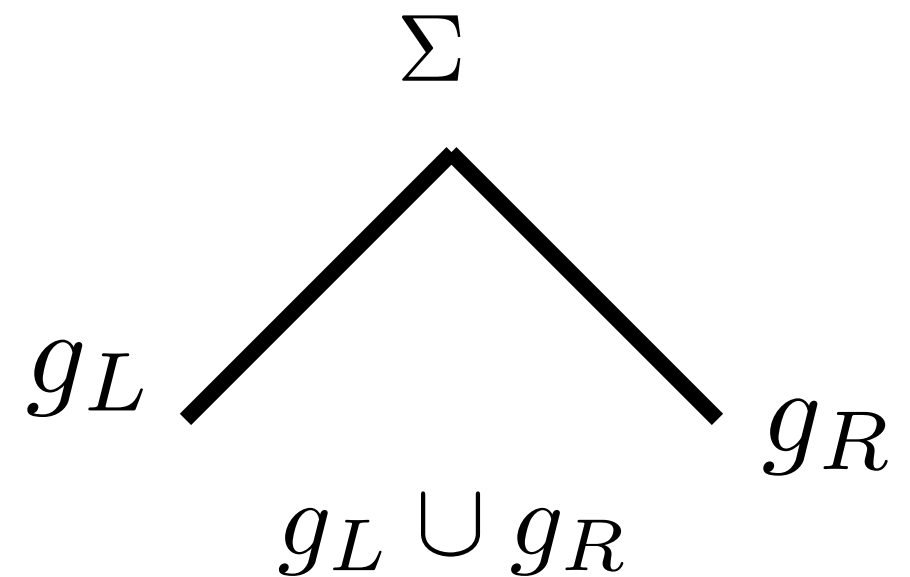
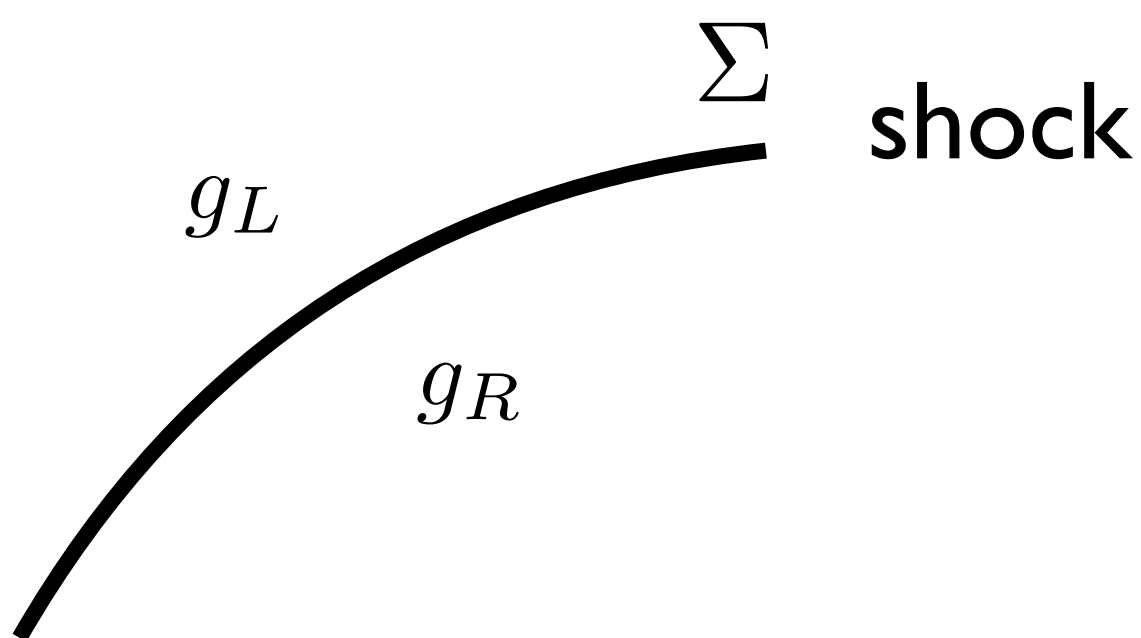
Given $\Gamma, d\Gamma \in W^{m,p}$, does there exist a $W^{m+2,p}$ coordinate transformation $x \rightarrow y$ with Jacobian $J \in W^{m+1,p}$, such that in y -coordinates, $\tilde{\Gamma} \in W^{m+1,p}$??

Or for shock-waves:

Given $\Gamma, d\Gamma \in L^\infty$, does there exist a $C^{1,1}$ coordinate transformation $x \rightarrow y$ with Jacobian $J \in C^{0,1}$, such that in y -coordinates, $\tilde{\Gamma} \in C^{0,1}$??

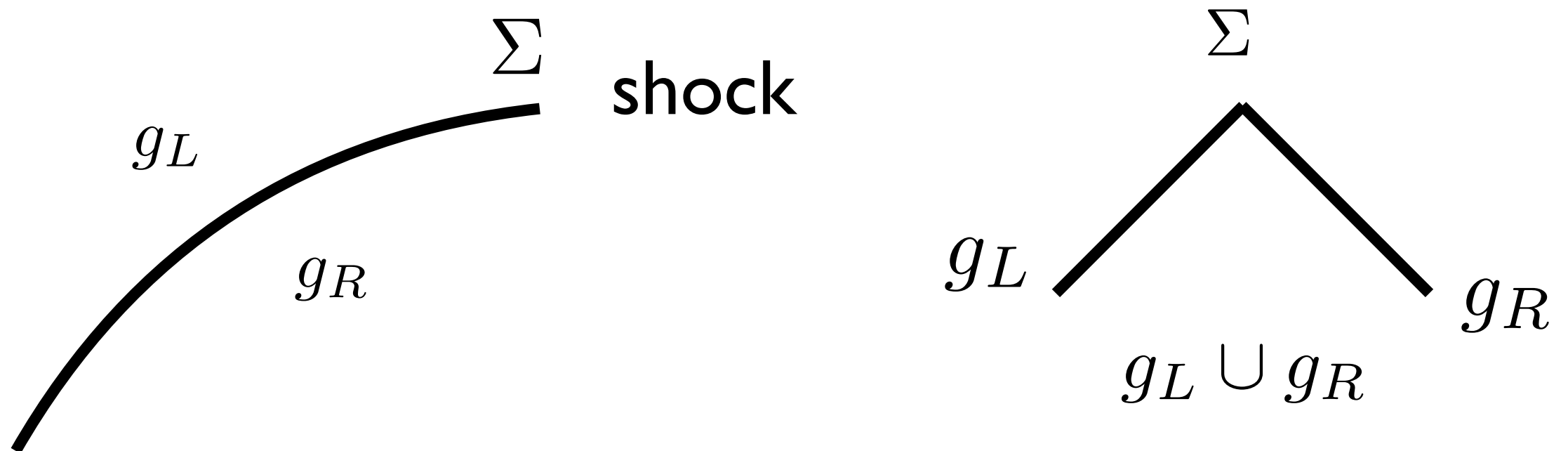
A result by Israel from the 60s resolves the issue for smooth shock surfaces in General Relativity:

Assume g_L and g_R are smooth solutions of the Einstein equations which match Lipschitz continuously across a smooth shock surface Σ , and let... $g \equiv g_L \cup g_R$



A result by Israel from the 60s resolves the issue for smooth shock surfaces in General Relativity:

Then $g \in C^{0,1}$, so delta functions exist in the second derivatives of the metric. The following theorem gives conditions under which they all cancel out in the curvature tensor...



Theorem (Israel/SmTe): The follow are equivalent:

- (1) $g = g_L \cup g_R$ is a weak solution of the Einstein equations with curvature in L^∞ .
- (2) All the delta functions cancel out in the Riemann curvature tensor.
- (3) The Second Fundamental Forms from g_L and g_R agree on the surface Σ
- (4) There exists a $C^{1,1}$ coordinate transformation in a neighborhood of Σ such that in the new coordinates,

$$g = g_L \cup g_R \in C^{1,1}$$

Theorem (Israel/SmTe):

Moreover, if any of the four equivalencies hold, then the Rankine-Hugoniot jump conditions (which express conservation of energy and momentum at the shock) also hold:

$$[G_{i\sigma}]n^\sigma = 0 = [T_{i\sigma}]n^\sigma$$

...on solutions of $G = \kappa T$

“Proof:” In Gaussian Normal Coordinates the components of the second fundamental form are $g_{ij,n}$ so if the second fundamental form is continuous, then the metric and its derivatives match on Σ , which implies $g \in C^{1,1}$ in GNC...

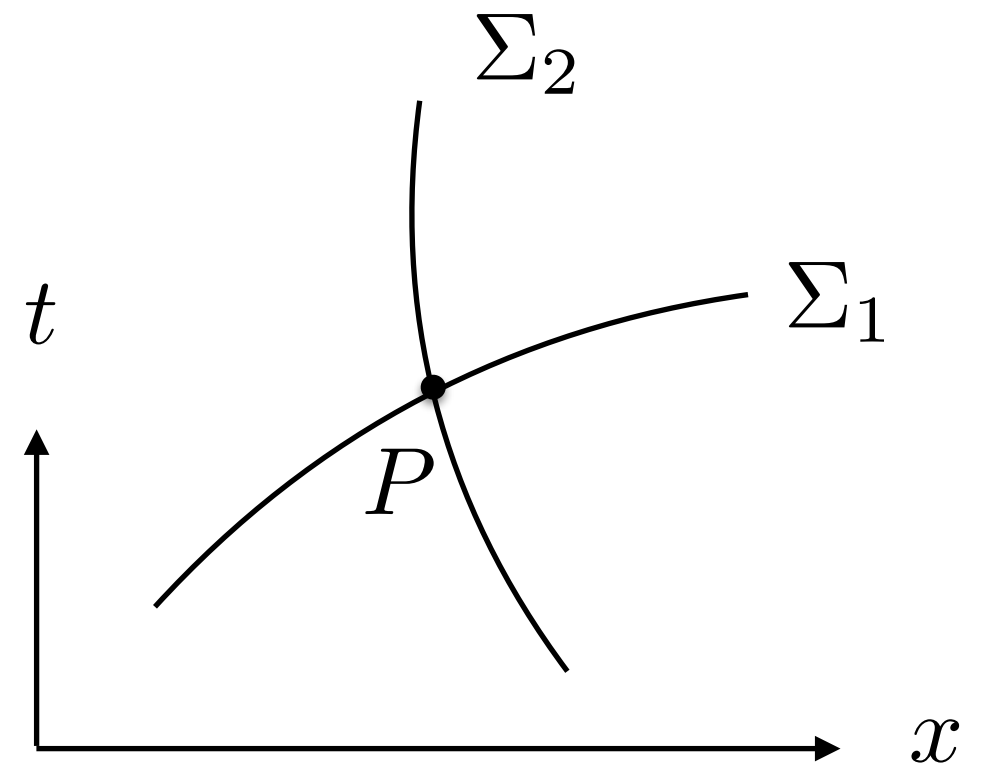
The map $x \rightarrow y$ to GNC is $C^{1,1}$...

...with Jacobian in $J \in C^{1,1}$. 

The first step forward from Israel's Theorem, was given by Moritz Reintjes:

Theorem (Reintjes): There always exists a $C^{1,1}$ transformation that smooths a $C^{0,1}$ metric to $C^{1,1}$ in a neighborhood of a point of regular shock-wave interaction in spherically symmetric spacetimes, between shock waves of different families...

Gaussian Normal Coordinates
break down at P:



Reintjes' procedure for finding the local coordinate systems of optimal smoothness is orders of magnitude more complicated than the Riemann normal, or Gaussian normal construction process.

The coordinate systems of optimal $C^{1,1}$ regularity are constructed by solving a complicated non-local PDE highly tuned to the structure of the interaction...

Trying to guess the coordinate system of optimal smoothness apriori, eg harmonic or Gaussian normal coordinates, didn't work.

Several **apparent miracles** happen in which the Rankine-Hugoniot jump conditions **come in to make seemingly over-determined equations consistent...**

... but...the principle behind what PDE's must be solved to smooth the metric in general, or when this is possible, appears entirely mysterious.

M. Reintjes, *Spacetime is Locally Inertial at Points of General Relativistic Shock Wave Interaction between Shocks from Different Characteristic Families*, Adv. Theor. Math. Phys., arXiv:1409.5060.

M. Reintjes and B. Temple, *No Regularity Singularities Exist at Points of General Relativistic Shock Wave Interaction between Shocks from Different Characteristic Families*,

Proc. R. Soc. A **471**:20140834.

<http://dx.doi.org/10.1098/rspa.2014.0834>

The first breakthrough in discovering the general principles at play in metric smoothing came with our discovery of the Riemann-flat condition...

The Riemann-flat Condition

The Riemann-flat Condition

“Shock Wave Interactions in General Relativity: The Geometry behind Metric Smoothing and the Existence of Locally Inertial Frames”

Moritz Reintjes, Blake Temple

<https://arxiv.org/abs/1610.02390>

The Riemann-flat condition:

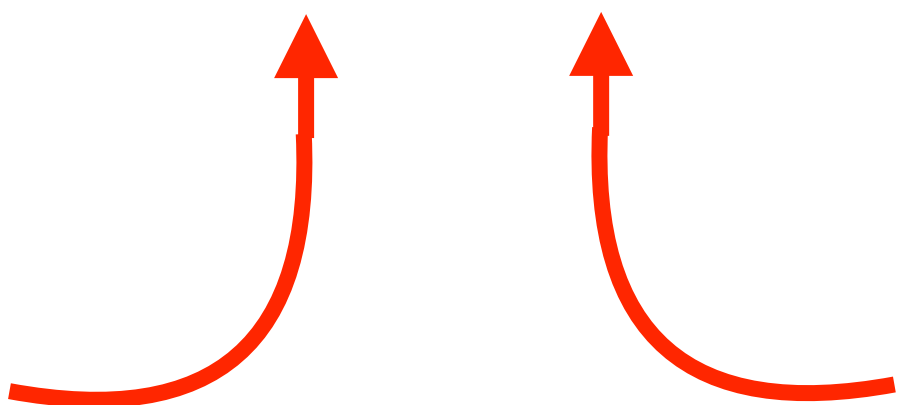
Assume $\Gamma, R \in L^\infty$.

Then: There exists a $C^{1,1}$ coordinate transformation that smooths an L^∞ connection Γ by one order to $C^{0,1}$ if and only if there exists a tensor $\tilde{\Gamma} \in C^{0,1}$ such that $Riem(\Gamma + \tilde{\Gamma}) = 0$.

In words: A smoothing transformation exists at shock-waves if and only if there exists a tensor, one order smoother than the original connection, such that when added to the original connection, the new connection is Riemann-flat.

In words: A smoothing transformation exists at shock-waves if and only if there exists a tensor, one order smoother than the original connection, such that when added to the original connection, the new connection is Riemann-flat.

$$Riem(\Gamma + \tilde{\Gamma}) = 0$$


$$\Gamma \in L^\infty \qquad \tilde{\Gamma} \in C^{0,1}$$

The same proof works at other orders of smoothness, for example: $\Gamma, \tilde{\Gamma} \in W^{m,p}$

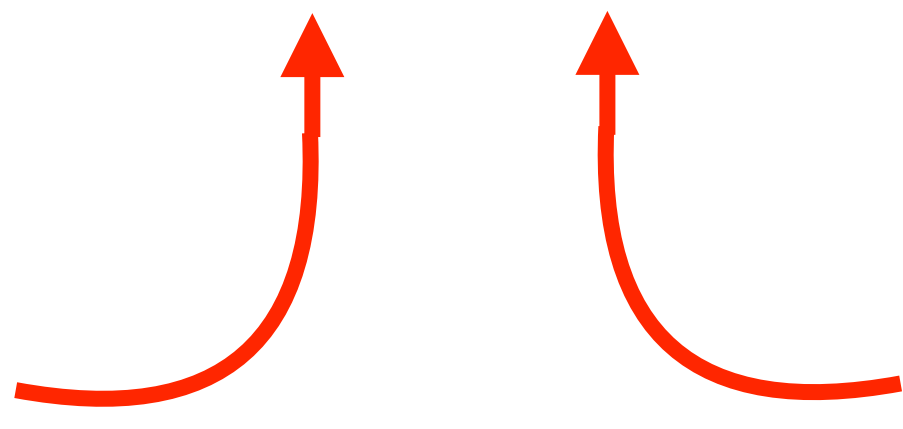
The same proof works at other orders of smoothness, for example: $\Gamma, \tilde{\Gamma} \in W^{m,p}$

A smoothing transformation $J \in W^{m+1,p}$ exists if and only if $\exists \tilde{\Gamma} \in W^{m+1,p}$ st

The same proof works at other orders of smoothness, for example: $\Gamma, \tilde{\Gamma} \in W^{m,p}$

A smoothing transformation $J \in W^{m+2,p}$ exists if and only if $\exists \tilde{\Gamma} \in W^{m+1,p}$ st

$$Riem(\Gamma + \tilde{\Gamma}) = 0$$



$\Gamma \in W^{m,p}$ $\Gamma \in W^{m+1,p}$

This shows that whether or not you can smooth the connection is a geometrical problem for connections, independent of the signature of the metric...

$$Riem(\Gamma + \tilde{\Gamma}) = 0$$

“Proof”: Assume $g \in C^{0,1}$, $\Gamma \in L^\infty$, $J \in C^{0,1}$

and

$$\Gamma_{\beta\gamma}^\alpha = \hat{\Gamma}_{jk}^i \frac{\partial y^\alpha}{\partial x^i} \frac{\partial x^j}{\partial y^\beta} \frac{\partial x^k}{\partial y^\gamma} + \underbrace{\frac{\partial^2 y^\alpha}{\partial x^\sigma \partial x^\tau} \frac{\partial x^\sigma}{\partial y^\beta} \frac{\partial x^\tau}{\partial y^\gamma}}_{\hat{\hat{\Gamma}}_{\beta\gamma}^\alpha \text{ Flat}}$$

L^∞ $C^{0,1}$ $C^{0,1}$

$-\tilde{\Gamma}_{\beta\gamma}^\alpha$

so $Riem(\Gamma + \tilde{\Gamma}) = 0$

The “hard” part is to show that if

$$Riem(\Gamma + \tilde{\Gamma}) = 0$$

...then a coordinate transformation exists which takes the connection $\Gamma + \tilde{\Gamma}$ to zero, and this is the coordinate transformation which smooths the original connection... ■

The Riemann-flat condition
reduces the problem of smoothing
Lorentzian spacetime metrics to
an equation on a tensor $\tilde{\Gamma}$ that
has all the remarkable properties
of the zero curvature condition of
Riemann.

Since $\tilde{\Gamma}$ continuous implies $\Gamma + \tilde{\Gamma}$ has the same jump discontinuities (shock set) as Γ at first we looked for a **Nash-type embedding theorem** for extending the shock set to a neighborhood as a flat connection in order to prove the metric can be smoothed.

However, our point of view changed again with another new idea in our next paper...

“The Regularity Transformation Equations: An elliptic mechanism for smoothing gravitational metrics in General Relativity”

Moritz Reintjes, Blake Temple

<https://arxiv.org/abs/1805.01004>

We set out to use the Riemann-flat condition for metric smoothing to derive a system of elliptic equations in unknowns $\tilde{\Gamma}$ and $J \dots$

The Regularity Transformation

Equations:

(RT-equations)

The RT-equations

“The Regularity Transformation Equations: An elliptic mechanism for smoothing gravitational metrics in General Relativity”

Moritz Reintjes, Blake Temple

<https://arxiv.org/abs/1805.01004>

The RT-equations:

$$\Delta \tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1}dJ) + d(J^{-1}A), \quad (1)$$

$$\Delta J = \delta(J\Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A, \quad (2)$$

$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle}), \quad (3)$$

$$\delta \vec{A} = v, \quad (4)$$

$$\text{Curl}(J) \equiv \partial_j J_i^\mu - \partial_i J_j^\mu = 0 \quad \text{on } \partial\Omega, \quad (5)$$

Here: $\tilde{\Gamma}$ is a matrix valued 1-form, J and A are matrix valued 0-forms, and \vec{J}, \vec{A} are vector valued 1-forms as follows:

$$\tilde{\Gamma} \equiv \tilde{\Gamma}_{\nu i}^{\mu} dx^i$$

$$J \equiv J_{\nu}^{\mu} \quad \vec{J} \equiv J_i^{\mu} dx^i \quad dJ = \text{Curl}(J)$$

$$A \equiv A_{\nu}^{\mu} \quad \vec{A} \equiv A_i^{\mu} dx^i \quad dA = \text{Curl}(A)$$

The integrability condition for J is: $\text{Curl}(J) = 0$

We introduce two new operations on matrix valued differential forms:

$$\overrightarrow{\text{div}}(\omega)^\alpha \equiv \sum_{l=1}^n \partial_l \left((\omega_l^\alpha)_{i_1, \dots, i_k} \right) dx^{i_1} \wedge \dots \wedge dx^{i_k}$$

(“take divergence in each component to create a vector valued form out of vector valued form”)

$$\langle A ; B \rangle_\nu^\mu \equiv \sum_{i_1 < \dots < i_k} A_{\sigma \ i_1 \dots i_k}^\mu B_{\nu \ i_1 \dots i_k}^\sigma$$

(“creates a matrix valued 0-form out of the “inner product” of two matrix valued k-forms”)

The RT-equations:

$$\Delta \tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1}dJ) + d(J^{-1}A), \quad (1)$$

$$\Delta J = \delta(J\Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A, \quad (2)$$

$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle}), \quad (3)$$

$$\delta \vec{A} = v, \quad (4)$$

$$\text{Curl}(J) \equiv \partial_j J_i^\mu - \partial_i J_j^\mu = 0 \quad \text{on } \partial\Omega, \quad (5)$$

The RT-equations:

$$\Delta \tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1}dJ) + d(J^{-1}A), \quad (1)$$

$$\Delta J = \delta(J\Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A, \quad (2)$$

$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle}), \quad (3)$$

$$\delta \vec{A} = v, \quad (4)$$

free to be chosen

$$\text{Curl}(J) \equiv \partial_j J_i^\mu - \partial_i J_j^\mu = 0 \quad \text{on } \partial\Omega, \quad (5)$$

The gauge freedom in the RT-equations is the freedom to choose v , together with the freedom to choose the boundary conditions in the $\tilde{\Gamma}$ and A equations...

Theorem (RT): Assume Γ is defined in a fixed coordinate system x on $\Omega \subset \mathbb{R}^n$, and $\Gamma, d\Gamma \in W^{m,p}(\Omega)$, $m \geq 1$, $p > n$.

If there exists $J \in W^{m+1,p}(\Omega)$ invertible, $\Gamma \in W^{m+1,p}(\Omega)$, $A \in W^{m,p}(\Omega)$ which solve the RT-equations,

Theorem (RT): Assume Γ is defined in a fixed coordinate system x on $\Omega \subset \mathbb{R}^n$, and $\Gamma, d\Gamma \in W^{m,p}(\Omega)$, $m \geq 1$, $p > n$.

If there exists $J \in W^{m+1,p}(\Omega)$ invertible, $\Gamma \in W^{m+1,p}(\Omega)$, $A \in W^{m,p}(\Omega)$ which solve the RT-equations,

Then

$$\tilde{\Gamma}' = \Gamma + J^{-1}dJ$$

solves the Riemann-flat condition.

Theorem (RT): Assume Γ is defined in a fixed coordinate system x on $\Omega \subset \mathbb{R}^n$, and $\Gamma, d\Gamma \in W^{m,p}(\Omega)$, $m \geq 1$, $p > n$.

If there exists $J \in W^{m+1,p}(\Omega)$ invertible, $\tilde{\Gamma} \in W^{m+1,p}(\Omega)$, $A \in W^{m,p}(\Omega)$ which solve the RT-equations,

then there exists a neighborhood $\Omega' \subset \Omega$ of p such that J is the Jacobian of a coordinate transformation $x \mapsto y$ on Ω' , and the components of Γ in y -coordinates are in $W^{m+1,p}(\Omega')$.

Theorem (RT): The converse also holds:

If there exists a coordinate transformation which smooths the connection,

Then the Jacobian together with the smoothed out connection solve the RT-equations for some A .

Conclude: The existence of a coordinate transformation which smooths a non-optimal connection by one order reduces to proving an existence theorem for the RT-equations, with

$\tilde{\Gamma}$ one order smoother than Γ

Existence for the RT-equations

“Optimal metric regularity in General Relativity follows from the RT-equations by elliptic regularity theory in L_p -spaces”

Moritz Reintjes, Blake Temple

<https://arxiv.org/abs/1808.06455>

Our New Existence Theorem:

Theorem (RT): Assume $\Gamma, d\Gamma \in W^{m,p}(\Omega)$ for $m \geq 1, p > n \geq 2$ in some coordinate system x . Then for each $q \in \Omega$ there exists a solution $(\tilde{\Gamma}, J, A)$ of the RT-equations defined in a neighborhood Ω_q of q such that

$$\tilde{\Gamma} \in W^{m+1,p}(\Omega_q), \quad J \in W^{m+1,p}(\Omega_q), \quad A \in W^{m,p}(\Omega_q).$$

Main Steps in the Derivation of the RT-equations

- Start with the Riemann-flat condition:

$$Riem(\Gamma + \tilde{\Gamma}) = 0$$

$$d\tilde{\Gamma} = d\Gamma + (\Gamma - \tilde{\Gamma}) \wedge (\Gamma - \tilde{\Gamma})$$

This can be viewed as an equation for $d\tilde{\Gamma}$

- Augment to a first order
Cauchy-Riemann system...

$$d\tilde{\Gamma} = d\Gamma + (\Gamma - \tilde{\Gamma}) \wedge (\Gamma - \tilde{\Gamma})$$

$$\delta\tilde{\Gamma} = h$$

But...this is not a solvable system

- To obtain a solvable system, we look to couple this Cauchy-Riemann system to an equation for the unknown Jacobian J .

Recall: $g \in C^{0,1}, \Gamma \in L^\infty, J \in C^{0,1}$

and

$$\Gamma_{\beta\gamma}^\alpha = \hat{\Gamma}_{jk}^i \frac{\partial y^\alpha}{\partial x^i} \frac{\partial x^j}{\partial y^\beta} \frac{\partial x^k}{\partial y^\gamma} + \underbrace{\frac{\partial^2 y^\alpha}{\partial x^\sigma \partial x^\tau} \frac{\partial x^\sigma}{\partial y^\beta} \frac{\partial x^\tau}{\partial y^\gamma}}_{\hat{\hat{\Gamma}}_{\beta\gamma}^\alpha \text{ Flat}}$$

L^∞ $C^{0,1}$ $C^{0,1}$

$-\tilde{\Gamma}_{\beta\gamma}^\alpha$

so $Riem(\Gamma + \tilde{\Gamma}) = 0$

Recall: $g \in C^{0,1}, \Gamma \in L^\infty, J \in C^{0,1}$

and

$$\Gamma_{\beta\gamma}^\alpha = \hat{\Gamma}_{jk}^i \frac{\partial y^\alpha}{\partial x^i} \frac{\partial x^j}{\partial y^\beta} \frac{\partial x^k}{\partial y^\gamma} + \underbrace{\frac{\partial^2 y^\alpha}{\partial x^\sigma \partial x^\tau} \frac{\partial x^\sigma}{\partial y^\beta} \frac{\partial x^\tau}{\partial y^\gamma}}_{\hat{\hat{\Gamma}}_{\beta\gamma}^\alpha \text{ Flat}}$$

L^∞ (pointing to $\Gamma_{\beta\gamma}^\alpha$)
 $C^{0,1}$ (pointing to $\hat{\Gamma}_{jk}^i$)
 $C^{0,1}$ (pointing to $\frac{\partial x^j}{\partial y^\beta}$ and $\frac{\partial x^k}{\partial y^\gamma}$)

$-\tilde{\Gamma}_{\beta\gamma}^\alpha$ (under the first term)
 $\hat{\hat{\Gamma}}_{\beta\gamma}^\alpha$ Flat (under the second term)

so $Riem(\Gamma + \tilde{\Gamma}) = 0$ or $\Gamma - \tilde{\Gamma} = J^{-1}dJ$

- To obtain a solvable system, we look to couple this Cauchy-Riemann system to an equation for the unknown Jacobian J .

Theorem: The Riemann-flat condition is equivalent to...

$$J^{-1}dJ = \Gamma - \tilde{\Gamma}$$

“Proof”...

Lemma: The following identity holds:

$$d(J^{-1}dJ) = J^{-1}dJ \wedge J^{-1}dJ$$

So assume $(J^{-1}dJ) = \Gamma - \tilde{\Gamma}$.

Then $d(J^{-1}dJ) = d\Gamma - d\tilde{\Gamma}$

which implies the Riemann-flat condition

$$d\tilde{\Gamma} = d\Gamma + (\Gamma - \tilde{\Gamma}) \wedge (\Gamma - \tilde{\Gamma})$$

- Thus we try to construct a closed system in $(\tilde{\Gamma}, J)$ out of two equivalent forms of the Riemann-flat condition...

$$d\tilde{\Gamma} = d\Gamma + (\Gamma - \tilde{\Gamma}) \wedge (\Gamma - \tilde{\Gamma})$$

$$\delta\tilde{\Gamma} = h$$

$$dJ = J(\Gamma - \tilde{\Gamma})$$

(They start out as equivalent!!)

- Thus we try to construct a closed system in $(\tilde{\Gamma}, J)$ out of two equivalent forms of the Riemann-flat condition...

$$d\tilde{\Gamma} = d\Gamma + (\Gamma - \tilde{\Gamma}) \wedge (\Gamma - \tilde{\Gamma})$$

$$\delta\tilde{\Gamma} = h$$

$$dJ = J(\Gamma - \tilde{\Gamma}) \qquad \delta J = 0$$

(for 0-forms)

(They start out as equivalent!!)

- We next employ the identity

$$\Delta = d\delta + \delta d$$

to derive two semi-linear elliptic
Poisson equations, one for $\Delta\tilde{\Gamma}$
and one for ΔJ

- **Apply** $\Delta = d\delta + \delta d$ **to**

$$d\tilde{\Gamma} = d\Gamma + (\Gamma - \tilde{\Gamma}) \wedge (\Gamma - \tilde{\Gamma})$$

$$\delta\tilde{\Gamma} = h$$

$$dJ = J(\Gamma - \tilde{\Gamma})$$

- **Apply** $\Delta = d\delta + \delta d$ **to**

$$d\tilde{\Gamma} = d\Gamma + (\Gamma - \tilde{\Gamma}) \wedge (\Gamma - \tilde{\Gamma})$$

$$\delta\tilde{\Gamma} = h$$

$$dJ = J(\Gamma - \tilde{\Gamma})$$

$$\delta J = 0 \quad (\text{delta of a 0-form is 0!})$$

- **Apply** $\Delta = d\delta + \delta d$ **to**

$$d\tilde{\Gamma} = d\Gamma + (\Gamma - \tilde{\Gamma}) \wedge (\Gamma - \tilde{\Gamma})$$

$$\delta\tilde{\Gamma} = h$$

$$dJ = J(\Gamma - \tilde{\Gamma})$$

$$\delta J = 0 \quad (\text{delta of a 0-form is 0!})$$

...to obtain

...to obtain

$$\Delta\tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1}dJ) + d(J^{-1}A),$$

$$\Delta J = \delta(J\Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A,$$

where $A = Jh$ is free...

- To impose the integrability condition

$$Curl(\vec{J}) = 0$$

...we require that **d** of the **vectorized**
right hand side of the J -**equation** **vanish**

$$\Delta J = \delta(J \Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A,$$

which gives the A -equation

$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle})$$

This leads to the RT-equations:

$$\Delta \tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1}dJ) + d(J^{-1}A), \quad (1)$$

$$\Delta J = \delta(J\Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A, \quad (2)$$

$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle}), \quad (3)$$

$$\delta \vec{A} = v, \quad (4)$$

$$\text{Curl}(J) \equiv \partial_j J_i^\mu - \partial_i J_j^\mu = 0 \quad \text{on } \partial\Omega, \quad (5)$$

- What I haven't shown you is how terms involving $\delta\Gamma$ which initially appear to be one derivative too low on the RHS, can be replaced by terms involving $d\Gamma$

To make the RHS smooth enough so that Δ formally lifts $\tilde{\Gamma}$ to one derivative above Γ

This property comes about by a
rather miraculous identity...

The RT-equations:

$$\Delta \tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1}dJ) + d(J^{-1}A), \quad (1)$$

$$\Delta J = \delta(J\Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A, \quad (2)$$

$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle}), \quad (3)$$

$$\delta \vec{A} = v, \quad (4)$$

$$\text{Curl}(J) \equiv \partial_j J_i^\mu - \partial_i J_j^\mu = 0 \quad \text{on } \partial\Omega, \quad (5)$$

Consider the A-equation:

$$\Delta \tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1} dJ) + d(J^{-1} A), \quad (1)$$

$$\Delta J = \delta(J \Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A, \quad (2)$$

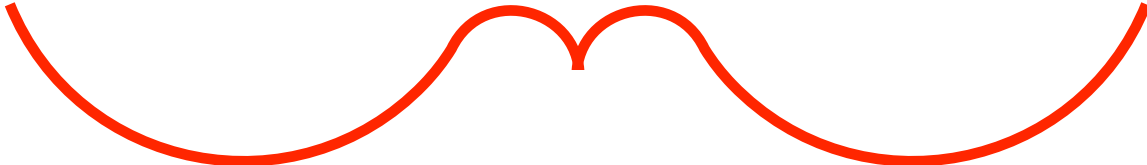
$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle}), \quad (3)$$

Consider the A-equation:

$$\Delta \tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1} dJ) + d(J^{-1} A), \quad (1)$$

$$\Delta J = \delta(J \Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A, \quad (2)$$

$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle}), \quad (3)$$



$$d(\overrightarrow{\delta(J \Gamma)})$$

Lemma:

Let $\Gamma \in W^{m,p}(\Omega)$ and $J \in W^{m+1,p}(\Omega)$ for $p > n$ and $m \geq 1$, then

$$d(\overrightarrow{\delta(J\Gamma)}) = \overrightarrow{\operatorname{div}}(dJ \wedge \Gamma) + \overrightarrow{\operatorname{div}}(J d\Gamma)$$

$d\delta \Gamma \in W^{-1,p}$ $\nabla(d\Gamma) \in W^{0,p}$

$d\Gamma$ **one derivative smoother than $\delta\Gamma$!**

- Finally, the $\tilde{\Gamma}$ which solve the RT-equations may not solve the Riemann-flat condition...

- Finally, the $\tilde{\Gamma}$ which solve the RT-equations may not solve the Riemann-flat condition...

Theorem: If $(\tilde{\Gamma}, J)$ solves the RT-equations, then

$$\tilde{\Gamma}' = \Gamma - J^{-1}dJ$$

- Finally, the $\tilde{\Gamma}$ which solve the RT-equations may not solve the Riemann-flat condition...

Theorem: If $(\tilde{\Gamma}, J)$ solves the RT-equations, then

$$\tilde{\Gamma}' = \Gamma - J^{-1}dJ$$

is Riemann-flat...

$$Riem(\Gamma - \tilde{\Gamma}') = 0$$

- Finally, the $\tilde{\Gamma}$ which solve the RT-equations may not solve the Riemann-flat condition...

Theorem: If $(\tilde{\Gamma}, J)$ solves the RT-equations, then

$$\tilde{\Gamma}' = \Gamma - J^{-1}dJ$$

is Riemann-flat...

$$Riem(\Gamma - \tilde{\Gamma}') = 0$$

And... $\tilde{\Gamma}'$ has the same regularity as Γ

- Finally, the $\tilde{\Gamma}$ which solve the RT-equations may not solve the Riemann-flat condition...

Theorem: If $(\tilde{\Gamma}, J)$ solves the RT-equations, then

$$\tilde{\Gamma}' = \Gamma - J^{-1}dJ$$

$\tilde{\Gamma}' \in W^{m+1,p}$ $\Gamma \in W^{m,p}$ $dJ \in W^{m,p}$

- Finally, the $\tilde{\Gamma}$ which solve the RT-equations may not solve the Riemann-flat condition...

Theorem: If $(\tilde{\Gamma}, J)$ solves the RT-equations, then

$$\tilde{\Gamma}' = \Gamma - J^{-1}dJ$$

$\tilde{\Gamma}' \in W^{m+1,p}$ $\Gamma \in W^{m,p}$ $dJ \in W^{m,p}$

$\tilde{\Gamma}'$ has the same regularity as $\tilde{\Gamma}$!

- Finally, the $\tilde{\Gamma}$ which solve the RT-equations may not solve the Riemann-flat condition...

Theorem: If $(\tilde{\Gamma}, J)$ solves the RT-equations, then

$$\tilde{\Gamma}' = \Gamma - J^{-1}dJ$$

$\tilde{\Gamma}' \in W^{m+1,p}$
 $\Gamma \in W^{m,p}$
 $dJ \in W^{m,p}$

The jumps in the $(m+1)$ -derivatives of $\tilde{\Gamma}'$ cancel out on the RHS!

- Finally, the $\tilde{\Gamma}$ which solve the RT-equations may not solve the Riemann-flat condition...

Theorem: If $(\tilde{\Gamma}, J)$ solves the RT-equations, then

$$\tilde{\Gamma}' = \tilde{\Gamma} - J^{-1}dJ$$

For fixed $J \in W^{m+1,p}$:

The transformation: $\tilde{\Gamma} \rightarrow \tilde{\Gamma}' \in W^{m+1,p}$

Represents a change of gauge: $A \rightarrow A' \in W^{m,p}$

Theorem: $(\tilde{\Gamma}, J, A)$ is a solution of the
RT-equations

Theorem: $(\tilde{\Gamma}, J, A)$ **is a solution of the RT-equations**

$$\Delta \tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1} dJ) + d(J^{-1} A), \quad (1)$$

$$\Delta J = \delta(J \Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A, \quad (2)$$

$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle}), \quad (3)$$

$$\delta \vec{A} = v, \quad (4)$$

$$\text{Curl}(J) \equiv \partial_j J_i^\mu - \partial_i J_j^\mu = 0 \quad \text{on } \partial\Omega,$$

Theorem: $(\tilde{\Gamma}, J, A)$ **is a solution of the RT-equations**

$$\Delta \tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1} dJ) + d(J^{-1} A), \quad (1)$$

$$\Delta J = \delta(J \Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A, \quad (2)$$

$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle}), \quad (3)$$

$$\delta \vec{A} = v, \quad (4)$$

$$Curl(J) \equiv \partial_j J_i^\mu - \partial_i J_j^\mu = 0 \quad \text{on } \partial\Omega,$$

if and only if $(\tilde{\Gamma}', J, A')$ **is a solution.**

- **Summary:** Starting with two equivalent forms of the RT-equations, we turn the first order equations into independent second order equations which allow for more general boundary conditions.

The second order equations don't imply the first order equations, but miraculously, a gauge transformation converts any solution into one which does satisfy the Riemann-flat condition.

Steps in the existence proof for the RT-equations

“Optimal metric regularity in General Relativity follows from the RT-equations by elliptic regularity theory in L_p -spaces”

Moritz Reintjes, Blake Temple

<https://arxiv.org/abs/1808.06455>

- The existence proof is based on an iteration scheme which applies the L^p theory of elliptic regularity at each stage

The L^p -theory of derivatives is a linear theory, and the RT-equations are nonlinear, so an iteration scheme is required...

The proof that the iterates converge
relies on only two theorems from
classical elliptic PDE theory...

Theorem (Elliptic Regularity): *Let $f \in W^{m-1,p}(\Omega)$, $m \geq 1$, and $u_0 \in W^{m+\frac{p-1}{p},p}(\partial\Omega)$ both be scalar functions. Assume $u \in W^{m+1,p}(\Omega)$ solves the Poisson equation $\Delta u = f$ with Dirichlet data $u|_{\partial\Omega} = u_0$. Then there exists a constant $C > 0$ depending only on Ω, m, n, p such that*

$$\|u\|_{W^{m+1,p}(\Omega)} \leq C \left(\|f\|_{W^{m-1,p}(\Omega)} + \|u_0\|_{W^{m+\frac{p-1}{p},p}(\partial\Omega)} \right).$$

Theorem (Elliptic Regularity): *Let $f \in W^{m-1,p}(\Omega)$, $m \geq 1$, and $u_0 \in W^{m+\frac{p-1}{p},p}(\partial\Omega)$ both be scalar functions. Assume $u \in W^{m+1,p}(\Omega)$ solves the Poisson equation $\Delta u = f$ with Dirichlet data $u|_{\partial\Omega} = u_0$. Then there exists a constant $C > 0$ depending only on Ω, m, n, p such that*

$$\|u\|_{W^{m+1,p}(\Omega)} \leq C \left(\|f\|_{W^{m-1,p}(\Omega)} + \|u_0\|_{W^{m+\frac{p-1}{p},p}(\partial\Omega)} \right).$$

Theorem (Gaffney Inequality): *Let $u \in W^{m+1,p}(\Omega)$ be a k -form for $m \geq 0$, $1 \leq k \leq n-1$ and (for simplicity) $n \geq 2$. Then there exists a constant $C > 0$ depending only on Ω, m, n, p , such that*

$$\|u\|_{W^{m+1,p}(\Omega)} \leq C \left(\|du\|_{W^{m,p}(\Omega)} + \|\delta u\|_{W^{m,p}(\Omega)} + \|u\|_{W^{m+\frac{p-1}{p},p}(\partial\Omega)} \right).$$

(Actually, what we found recorded:)

Theorem (Elliptic Regularity): *Let $u \in W^{2,p}(\Omega)$ be a scalar, $1 < p < \infty$. Then there exists a constant $C > 0$ depending only on Ω, m, n, p , such that*

$$\|u\|_{W^{2,p}(\Omega)} \leq C \left(\|\Delta u\|_{L^p(\Omega)} + \|u\|_{W^{1,p}(\Omega)} + \|u\|_{W^{2-\frac{1}{p},p}(\partial\Omega)} \right).$$

(What we need:)

Theorem (Elliptic Regularity): *Let $f \in W^{m-1,p}(\Omega)$, $m \geq 1$, and $u_0 \in W^{m+\frac{p-1}{p},p}(\partial\Omega)$ both be scalar functions. Assume $u \in W^{m+1,p}(\Omega)$ solves the Poisson equation $\Delta u = f$ with Dirichlet data $u|_{\partial\Omega} = u_0$. Then there exists a constant $C > 0$ depending only on Ω, m, n, p such that*

$$\|u\|_{W^{m+1,p}(\Omega)} \leq C \left(\|f\|_{W^{m-1,p}(\Omega)} + \|u_0\|_{W^{m+\frac{p-1}{p},p}(\partial\Omega)} \right).$$

References:

L. C. Evans, *Partial Differential Equations*, Berkeley Mathematics Lecture Notes, **3A**, 1994.

G. Csató, B. Dacorogna and O. Kneuss, *The Pullback Equation for Differential Forms*, Birkhäuser, Progress in Nonlinear Differential Equations and Their Applications, Vol. 83, ISBN: 978-0-8176-8312-2.

D. Gilbarg and N. Trudinger, *Elliptic Partial Differential Equations of Second Order*, 3rd edition, (1997), Springer Verlag, ISBN 3-540-41160-7.

References:

P. Grisvard, *Elliptic Problems in Nonsmooth Domains*, SIAM ed. (2011), ISBN 978-1-611972-02-3.

S. Agmon, A. Douglas, L. Nirenberg, *Estimates Near the Boundary for Solutions of Elliptic Partial Differential Equations Satisfying General Boundary Conditions*, Comm. Pure Appl. Math, Vol **12**, 623-727 (1959)

- One of the main obstacles to overcome was how to reduce the existence theorem at each iterate to a problem with Dirichlet boundary conditions... so standard linear elliptic regularity applies...

For this we introduce ancillary variable y so that $dy = \vec{J}$, $d^2y = d\vec{J} = \text{Curl}(J)$

- Details of Proof in Moritz's Talk!

Conclusion

- We show that **proving** non-optimal metrics (or connections) can be **smoothed** one order by coordinate transformation is **equivalent** to **proving existence** for the RT-equations...

- We show that **proving** non-optimal metrics (or connections) can be **smoothed** one order by coordinate transformation is **equivalent** to proving **existence** for the RT-equations...
- We prove existence for the RT-equations above the threshold smoothness of...

$$\Gamma, d\Gamma \in W^{m,p} \quad m \geq 1, \quad m > n$$

- As a Corollary we have that solutions constructed in SSC can be smoothed one order by coordinate transformation...

- As a Corollary we have that solutions constructed in SSC can be smoothed one order by coordinate transformation...

Q: What do the SSC metrics look like in coordinates of optimal regularity?

- As a Corollary we have that solutions constructed in SSC can be smoothed one order by coordinate transformation...

Q: What do the SSC metrics look like in coordinates of optimal regularity?

Ans: We don't know!

- Q: How does this fit in with other methods of obtaining solutions in GR?

- Q: How does this fit in with other methods of obtaining solutions in GR?

Existence Theorems on the IVP in GR employ apriori coordinate ansatz's, like Harmonic Coordinates...how does this fit in?

- Q: How does this fit in with other methods of obtaining solutions in GR?

Existence Theorems on the IVP in GR employ *a priori* coordinate ansatz's, like Harmonic Coordinates...how does this fit in?

Theorems on optimal regularity in GR begin with technical conditions on the spacetime (regularity of geodesic balls, etc) and mostly address vacuum spacetimes...Anderson, Kleinerman, Rodnianski, Dafermos, LeFloch...?

- Our theorem is geometric, applies independent of matter sources or metric signature, and makes no apriori assumptions on the spacetime other than its regularity...

- **Q:** Can the existence theory for the RT-equations extend to the case of **GR shock-waves** (topic of our current research)?

...the case $\Gamma, d\Gamma \in L^\infty$? or $m < 1$

- Main Obstacles for L^∞ :

- Main Obstacles for L^∞ :

The Laplacian doesn't lift derivatives two full orders when sources in L^∞

- Main Obstacles for L^∞ :

The Laplacian doesn't lift derivatives two full orders when sources in L^∞

...issue of Calderon-Zygmund Singularities?

- Main Obstacles for L^∞ :

The Laplacian doesn't lift derivatives two full orders when sources in L^∞

...issue of Calderon-Zygmund Singularities?

- Main Obstacles for L^∞ :

The Laplacian doesn't lift derivatives two full orders when sources in L^∞

...issue of Calderon-Zygmund Singularities?

Conclude: The RT-equations connect Regularity Singularities in GR to Calderon-Zygmund Singularities in elliptic PDE's...

- Main Obstacles for L^∞ :

The Laplacian doesn't lift derivatives two full orders when sources in L^∞

...issue of Calderon-Zygmund Singularities?

Conclude: The RT-equations connect Regularity Singularities in GR to Calderon-Zygmund Singularities in elliptic PDE's...

(Apparently two completely different kinds of singularities...?)

- Quadratic Nonlinearities on the RHS of RT-equations imply iterations may not stay in space...

L^∞ is closed under nonlinear products

L^p is NOT closed under nonlinear products

What about the space BMO?

- Quadratic Nonlinearities on the RHS of RT-equations imply iterations may not stay in space...

L^∞ is closed under nonlinear products

L^p is NOT closed under nonlinear products

What about the space BMO?

There is a natural distance from BMO to L^∞ ...?

- What is working for us in L^∞ ?

- What is working for us in L^∞ ?

Ans: Although the Laplacian does not lift all functions $f \in L^\infty$ up by two derivatives—i.e.,

$$\Delta u = f \in L^\infty$$

does not always imply $u \in C^{1,1} \dots$

- What is working for us in L^∞ ?

Ans: Although the Laplacian does not lift all functions $f \in L^\infty$ up by two derivatives—i.e.,

$$\Delta u = f \in L^\infty$$

does not always imply $u \in C^{1,1} \dots$

...there is an enormous gauge freedom on the RHS of the RT-equations.

To settle the problem of regularity singularities at GR shock waves, ... All we need is that there exists one solution of the RT-equations in $C^{1,1}$, for some v , and any choice of boundary conditions for $(\tilde{\Gamma}, A) \dots$

The RT-equations:

$$\Delta \tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1}dJ) + d(J^{-1}A), \quad (1)$$

$$\Delta J = \delta(J\Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A, \quad (2)$$

$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle}), \quad (3)$$

$$\delta \vec{A} = v, \quad (4)$$

free to be chosen

$$\text{Curl}(J) \equiv \partial_j J_i^\mu - \partial_i J_j^\mu = 0 \quad \text{on } \partial\Omega, \quad (5)$$

The RT-equations:

$$\Delta \tilde{\Gamma} = \delta d\Gamma - \delta d(J^{-1}dJ) + d(J^{-1}A), \quad (1)$$

$$\Delta J = \delta(J\Gamma) - \langle dJ; \tilde{\Gamma} \rangle - A, \quad (2)$$

$$d\vec{A} = \overrightarrow{\text{div}}(dJ \wedge \Gamma) + \overrightarrow{\text{div}}(J d\Gamma) - d(\overrightarrow{\langle dJ; \tilde{\Gamma} \rangle}), \quad (3)$$

$$\delta \vec{A} = v, \quad (4)$$

free to chose boundary conditions

$$\text{Curl}(J) \equiv \partial_j J_i^\mu - \partial_i J_j^\mu = 0 \quad \text{on } \partial\Omega, \quad (5)$$

Conclude:

If we can prove there exists solutions $\tilde{\Gamma} \in C^{1,1}$ of the RT-equations for $\Gamma, d\Gamma \in L^\infty$ then regularity singularities do not exist at GR shocks...

Conclude:

If we can prove there exists solutions $\tilde{\Gamma} \in C^{1,1}$ of the RT-equations for $\Gamma, d\Gamma \in L^\infty$ then regularity singularities do not exist at GR shocks...

...and it suffices to solve the Einstein equations in coordinates where the metric is non-optimal...

Conclude:

If we can prove there exists solutions $\tilde{\Gamma} \in C^{1,1}$ of the RT-equations for $\Gamma, d\Gamma \in L^\infty$ then regularity singularities do not exist at GR shocks...

...and it suffices to solve the Einstein equations in coordinates where the metric is non-optimal...

If we can prove there are no solutions of the RT-equations for $\Gamma, d\Gamma \in L^\infty$, then we have discovered a new kind of singularity in GR...

Conclude:

If we can prove there exists solutions $\tilde{\Gamma} \in C^{1,1}$ of the RT-equations for $\Gamma, d\Gamma \in L^\infty$ then regularity singularities do not exist at GR shocks...

...and it suffices to solve the Einstein equations in coordinates where the metric is non-optimal...

If we can prove there are no solutions of the RT-equations for $\Gamma, d\Gamma \in L^\infty$, then we have discovered a new kind of singularity in GR...

Either way its interesting!

References (Reintjes Temple):

[arXiv:1610.0239](#)

[arXiv:1808.06455](#)

[arXiv:1805.01004](#)

END

Thank

You!